

**AL-MG AGE OF THE ZAKLODZIE ENSTATITE METEORITE.** N. Sugiura<sup>1</sup> and W. Fujiya<sup>1</sup>, <sup>1</sup>Department of Earth and Planetary Science, University of Tokyo, Tokyo, Japan. E-mail [Sugiura@eps.s.u-tokyo.ac.jp](mailto:Sugiura@eps.s.u-tokyo.ac.jp).

**Introduction:** Zaklodzie is an ungrouped meteorite closely related to enstatite chondrites [1]. It is considered to be an achondrite but contains abundant metal and the bulk composition seems to be similar to enstatite chondrites [2]. Its texture suggests relatively rapid cooling and an impact melt origin was suggested [2]. However, mainly based on cumulate-like texture and enrichment of plagioclase, [3] suggested melting due to decay of <sup>26</sup>Al. Two generations of feldspar may be present because the feldspar compositions show large variations [3]. The U/Th-He age of 2.1 Ga and the K-Ar age of 4.4Ga were reported [4]. These gas retention ages are often much younger than the formation (initial crystallization) age. Here we report the Al-Mg age of this meteorite.

**Experimental:** A polished thin section was examined with a scanning electron microscope. Chemical compositions of feldspar grains were measured with energy dispersive spectroscopy. An ion probe (Cameca-6f) at the Univ. of Tokyo was used for the Al-Mg isotope measurements. Mg isotopes were measured with an electron multiplier whereas Al was measured with a Faraday cup in cases of feldspar. In the case of pyroxene measurements which were made for precise determination of y-intercept of the isochron diagram, Al was measured with an electron multiplier. It took about 25 minutes for one SIMS measurement. The measurement was made twice on the same spot.

**Results and discussion:** As reported in literatures [3], feldspar grains with a wide range of compositions were found. The Al/Mg ratios measured by SIMS are correlated with CaO (Fig.1), although the correlation is not very strong ( $r \sim 0.92$ ). Since CaO-poor feldspar was probably produced by a secondary event, some Mg seems to be added to CaO-poor feldspar together with Na. Al/<sup>24</sup>Mg ratios in feldspar range from  $\sim 1000$  (in CaO-poor feldspar) to  $\sim 5500$  (in CaO-rich feldspar). The high Al/<sup>24</sup>Mg ratios allowed relatively precise age determination. The  $\delta^{26}\text{Mg}$  values range from 0‰ to 13‰ and are correlated with the Al/<sup>24</sup>Mg ratios (Fig.2). This correlation is considered to be an isochron and the inferred initial <sup>26</sup>Al/<sup>27</sup>Al ratio is  $(3.1 \pm 1.1(2\sigma)) \times 10^{-7}$ . Relative to CAIs with the canonical value of <sup>26</sup>Al/<sup>27</sup>Al =  $5 \times 10^{-5}$ , the age of Zaklodzie is  $5.4 \pm 0.4(2\sigma)$  Ma.

It was noticed that the  $\delta^{26}\text{Mg}$  values for CaO-poor, second-generation feldspar were nearly normal. Within

the  $2\sigma$  error limit (Errors shown in Fig.2 are  $1\sigma$ .), they may be considered to belong to the isochron, or they may be considered to have been reset by a late secondary event. At present, the precision of the data is not good enough to distinguish these two alternative interpretations. The initial <sup>26</sup>Al/<sup>27</sup>Al ratio is hardly affected even if the data for the CaO-poor feldspar are omitted from the isochron.

If the absolute age of CAIs is taken as  $\sim 4567.1$  Ma [5], then the absolute age of Zaklodzie is about 4561.7 Ma. This is similar to Al-Mg ages of many basaltic achondrites. Many eucrites, quenched angrites, NWA 011 etc. have absolute ages of 4562~4563 Ma [5,6]. This old age suggests that the heat source for melting of Zaklodzie was likely to be decay of <sup>26</sup>Al rather than impact heating.

The relationship of Zaklodzie with enstatite achondrites has not been studied in detail. The Mn-Cr bulk-isochron age of enstatite achondrites is  $\sim 4563$  Ma [7] which is not very different from the age of Zaklodzie. But the bulk isochron has a high y-intercept which indicates that the source of enstatite achondrites evolved in an environment with an elevated Mn/Cr ratio for several million years [7]. Since the bulk composition of Zaklodzie seems to be nearly chondritic, it must have evolved in a different parentbody than the enstatite achondrites.

Old ages of ungrouped achondrites (Zaklodzie and NWA 011) suggest that there were many, tiny, molten planetesimals in the early solar nebula.

#### References:

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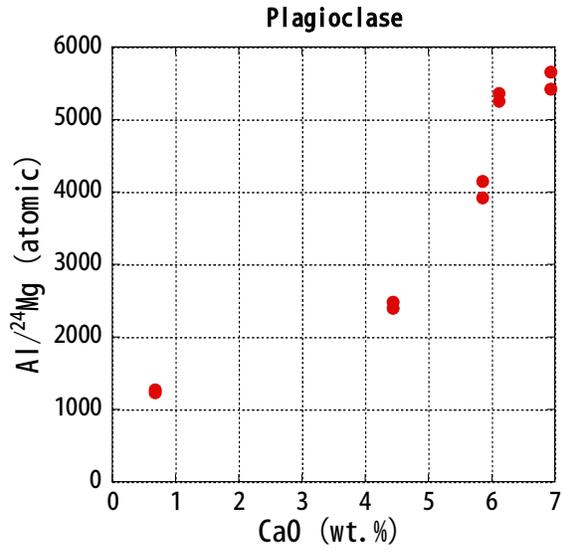


Fig.1 Al/<sup>24</sup>Mg ratios measured by SIMS are plotted against CaO contents measured by EDS.

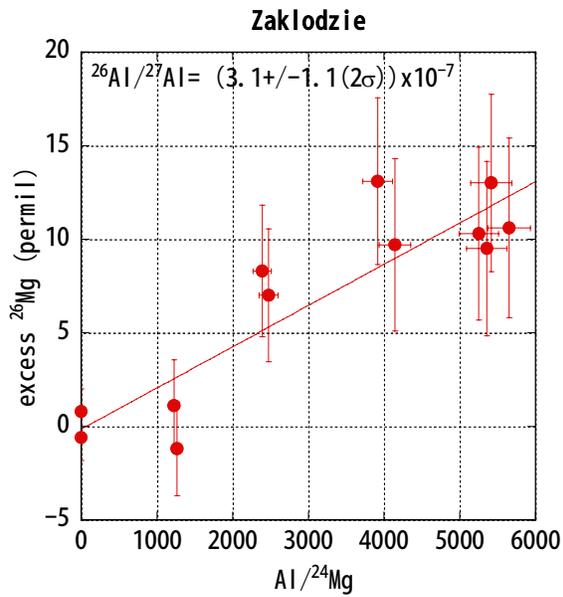


Fig.2 Al-Mg isochron for Zaklodzie.